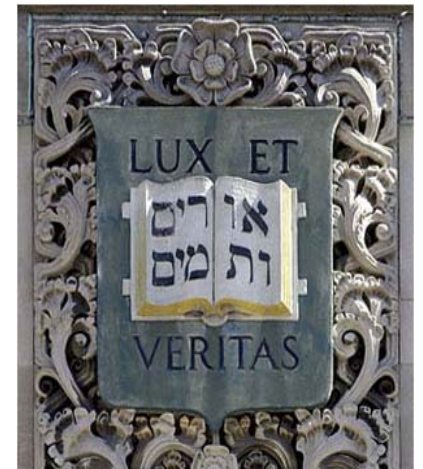

LIPSS results and status

O.K. Baker
For the LIPSS Collaboration

20-June-08
Fourth Patras Workshop

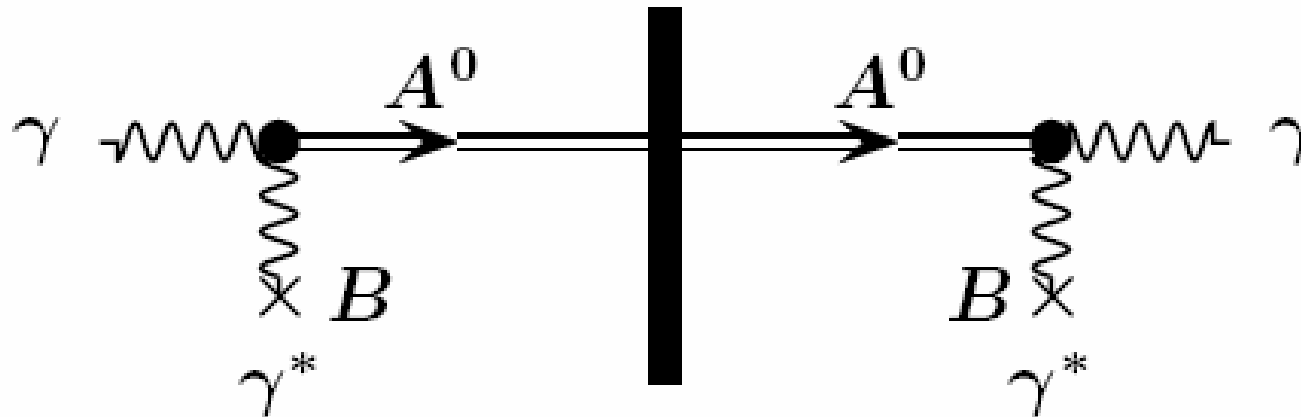


LIPSS at Jefferson Lab

- experimental (data) runs in Feb-March, 2007
 - submitted results for publication: see [arXiv:0806.2631](https://arxiv.org/abs/0806.2631)

- spent past year . . .
 - characterizing backgrounds better
 - analyzing data with improved methods
 - upgrading FEL optics and LIPSS experiment

photon regeneration
 'light shining through a wall'
www.desy.de/~ringwald
 Phys. Rev. D47 3707 (1993)



γ - A^0 (or A^0 - γ) conversion probability

$$P_{\gamma \rightarrow \phi} = \frac{1}{4} (gBL)^2 \left\{ \frac{\sin\left(\frac{m_\phi^2 L}{4\omega}\right)}{\frac{m_\phi^2 L}{4\omega}} \right\}^2$$

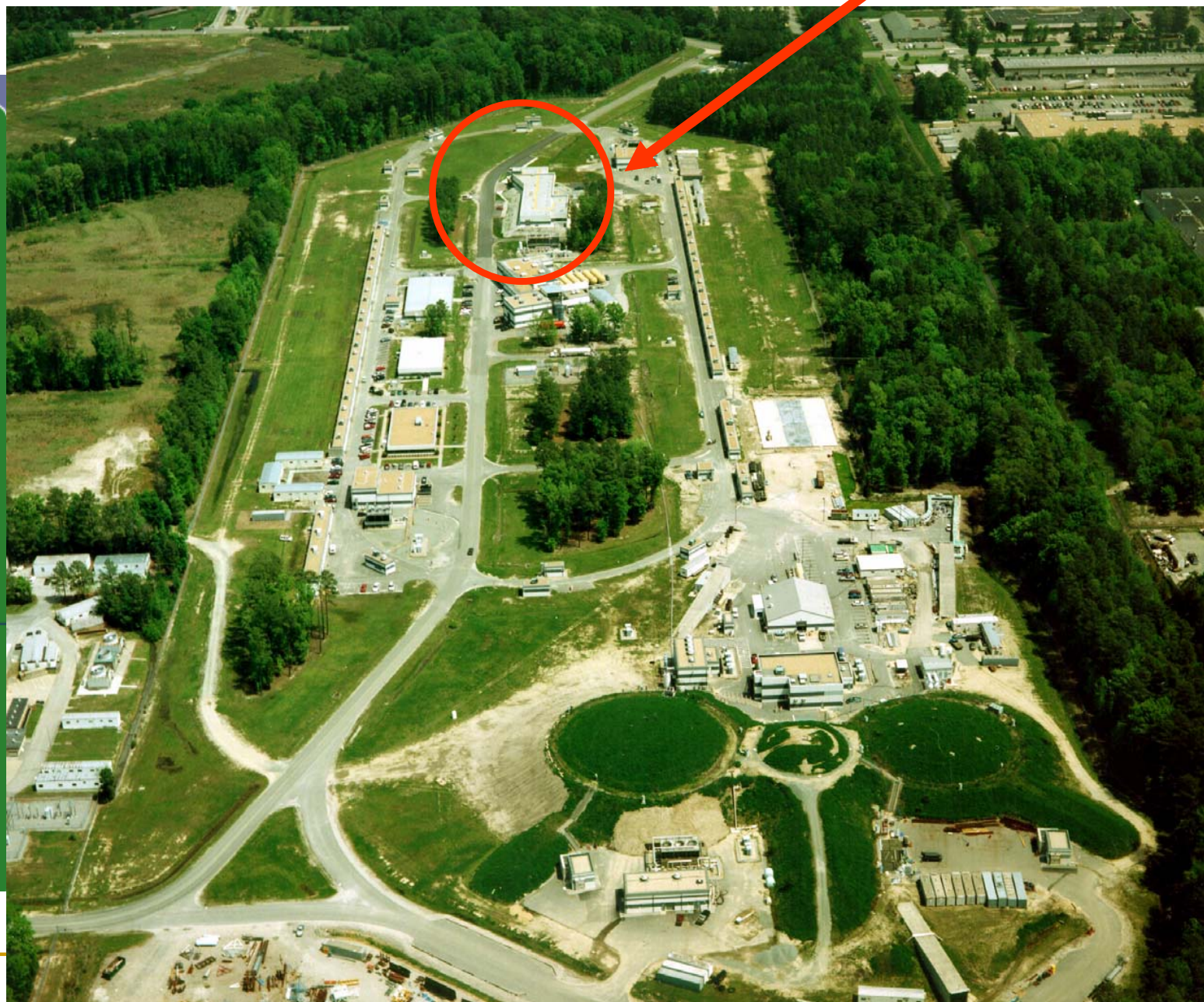
$$m_\phi^2 < 4\omega/L$$

- couple polarized laser light with magnetic field
- Sikivie (1983); Ansel'm (1985); Van Bibber et al (1987)

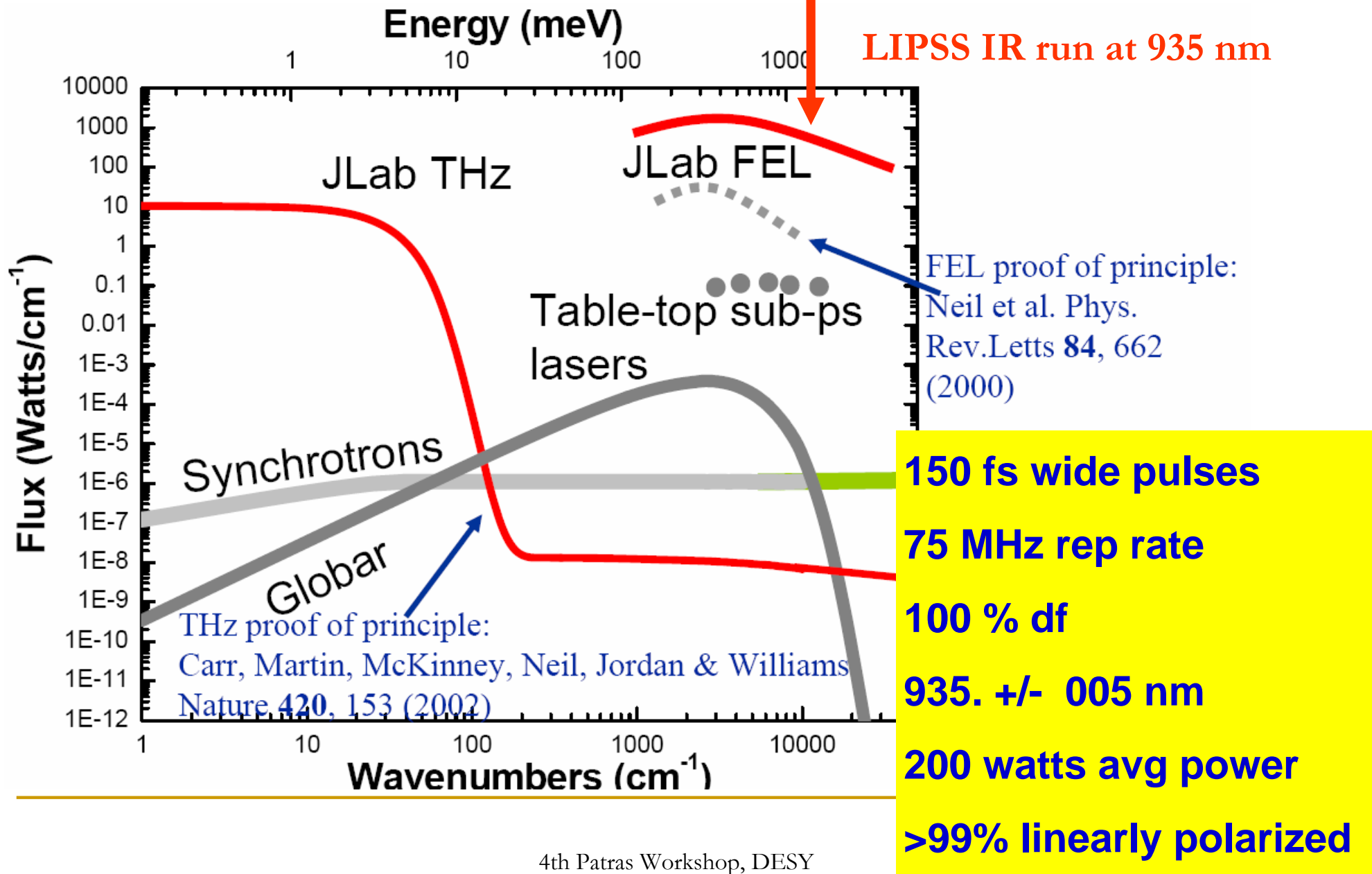
Jefferson Lab and the Free Electron Laser



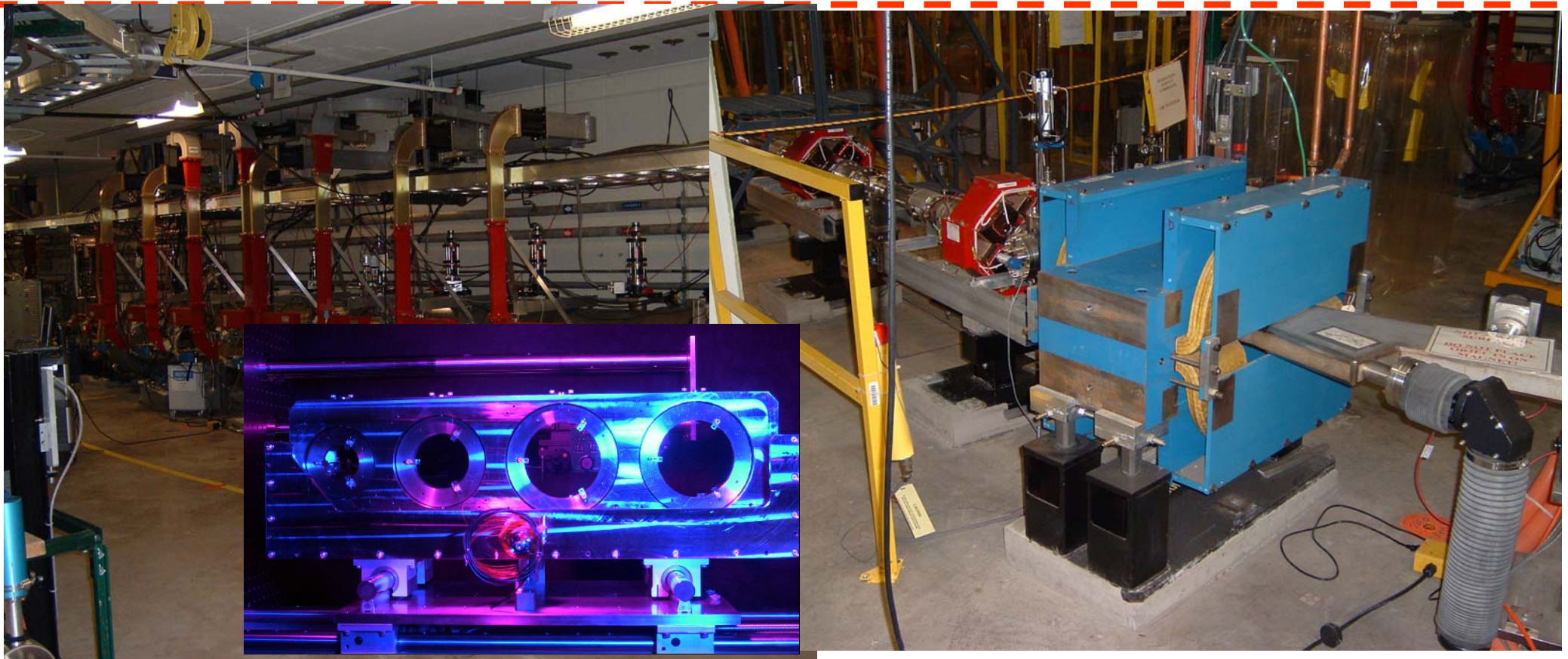
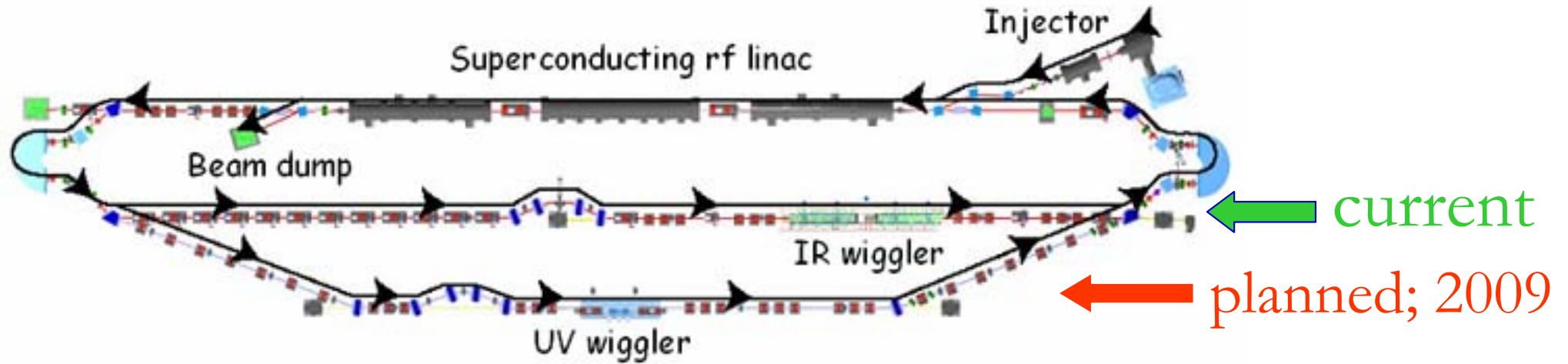
Jefferson Lab and the Free Electron Laser



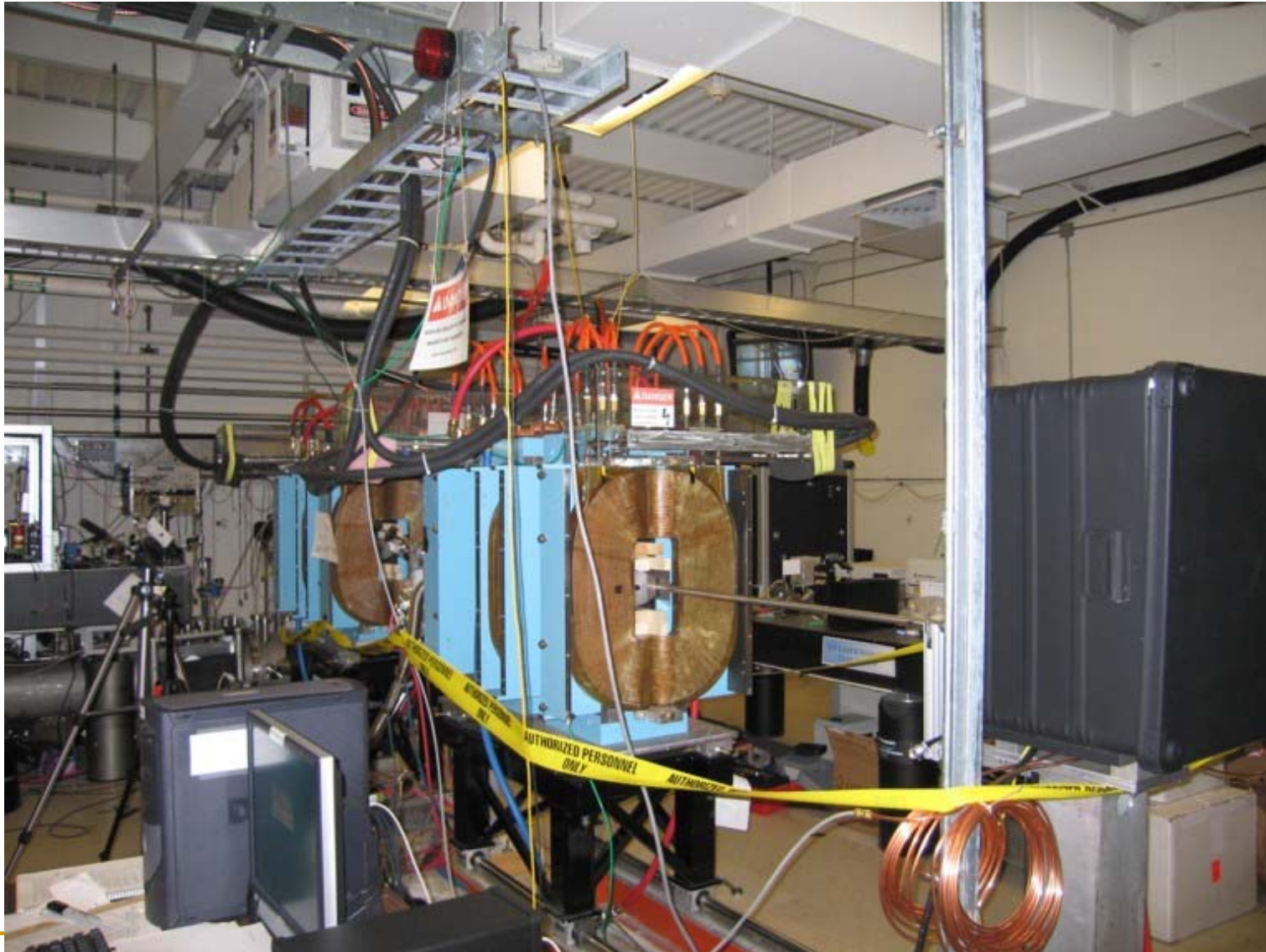
JLAB FEL spectroscopic range



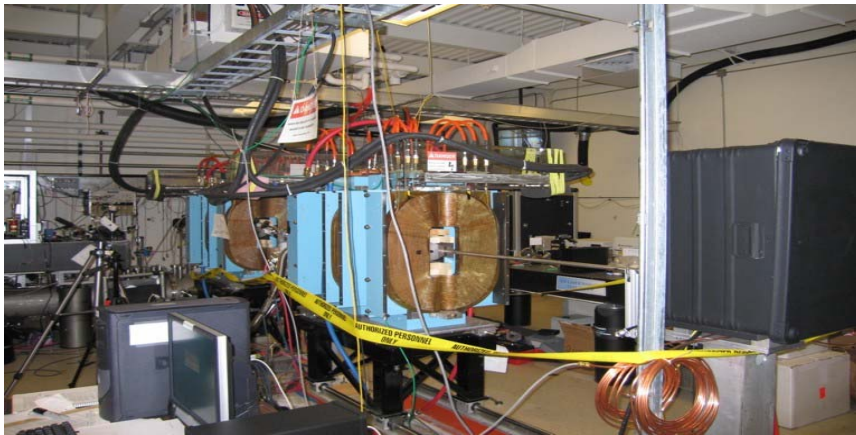
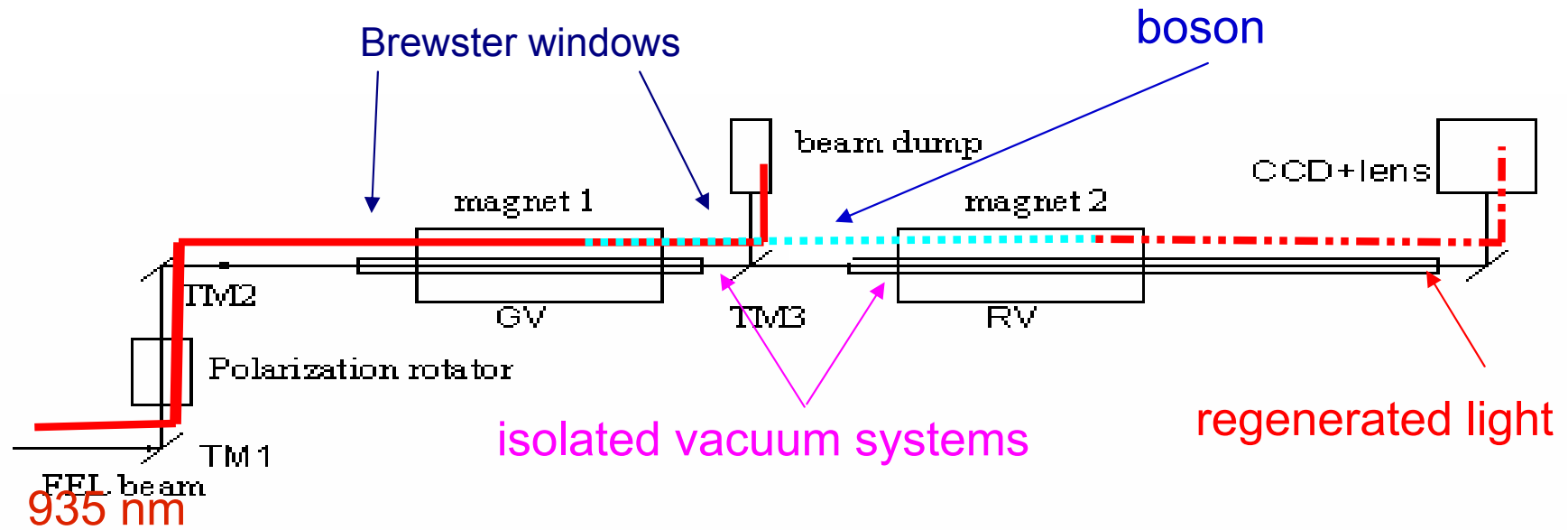
JLAB FEL: regeneration experiment

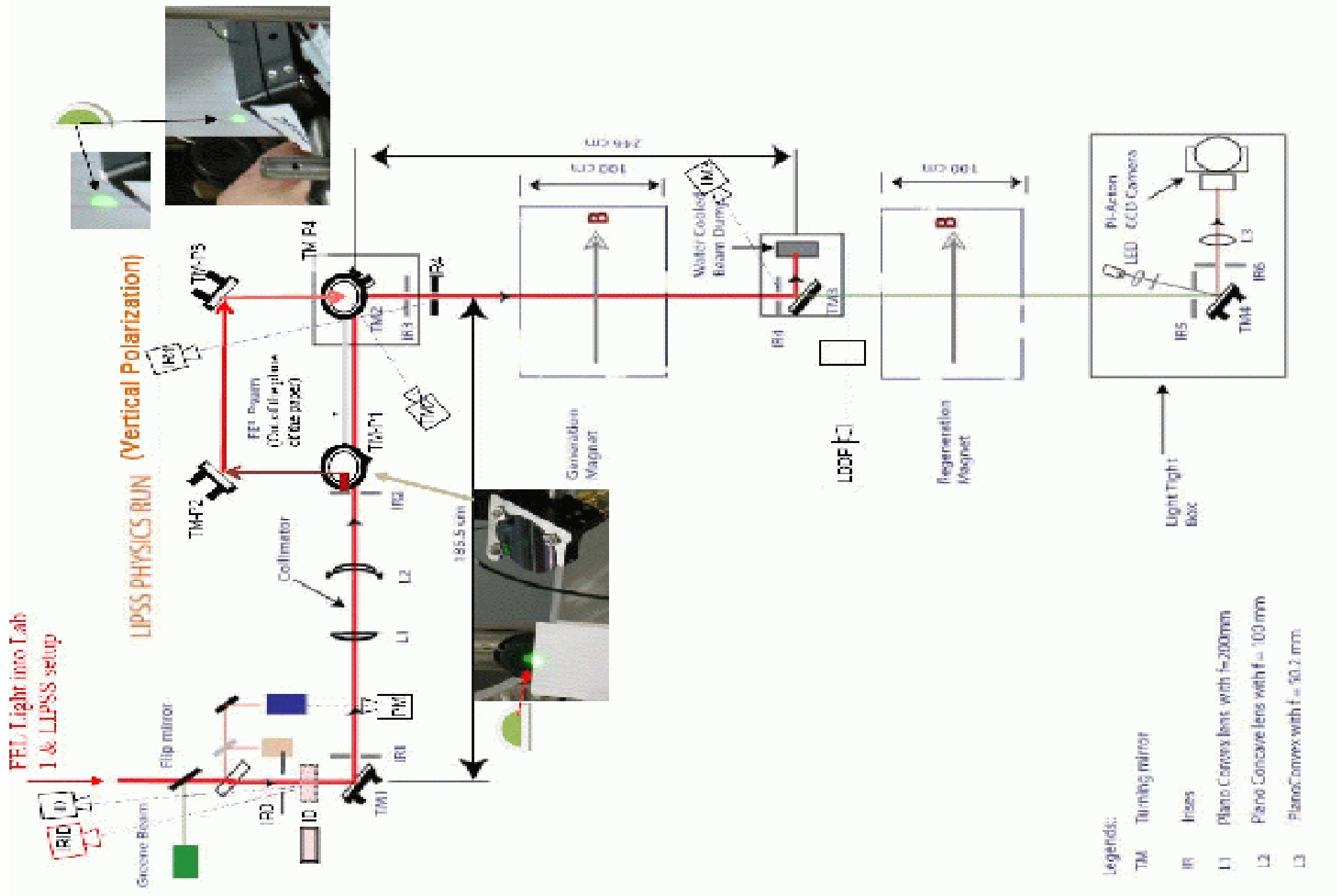


LIPSS at Jefferson Lab

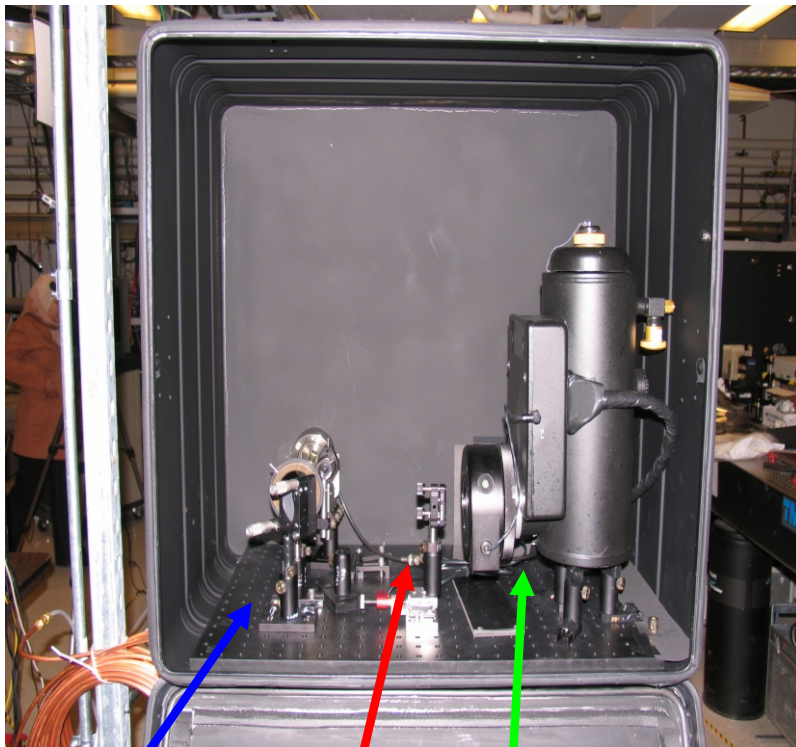


LIPSS – experiment schematic.



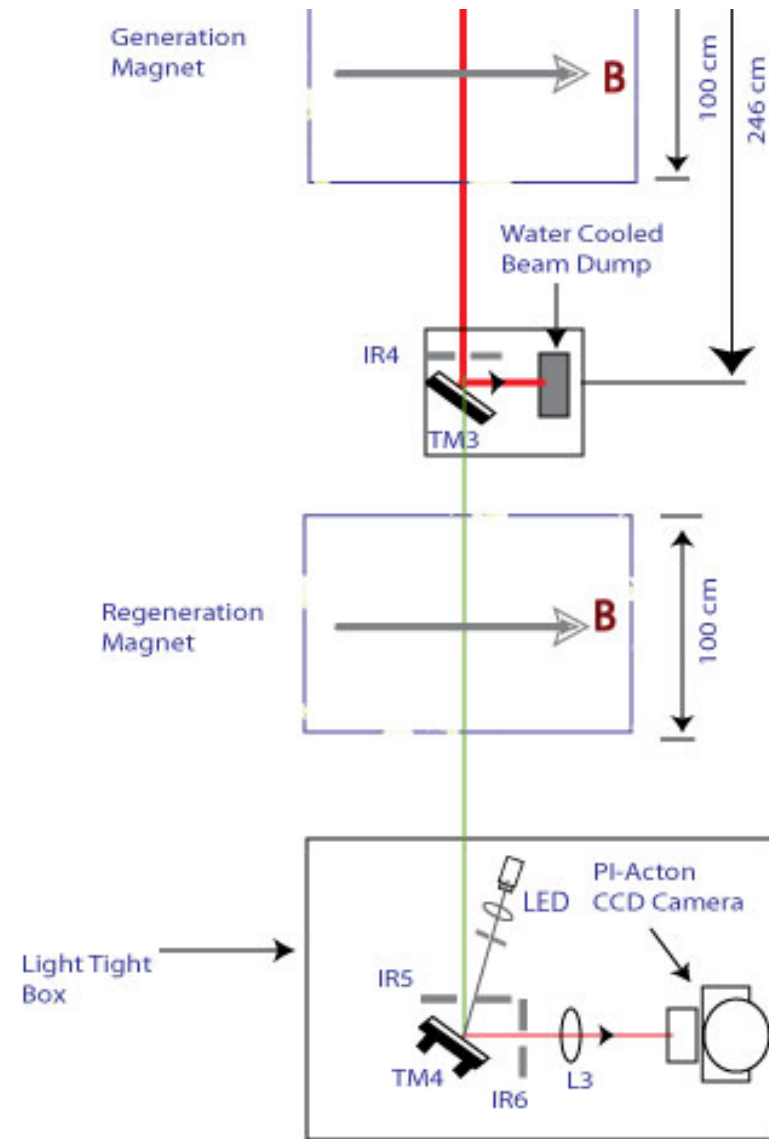


LIPSS detector chamber

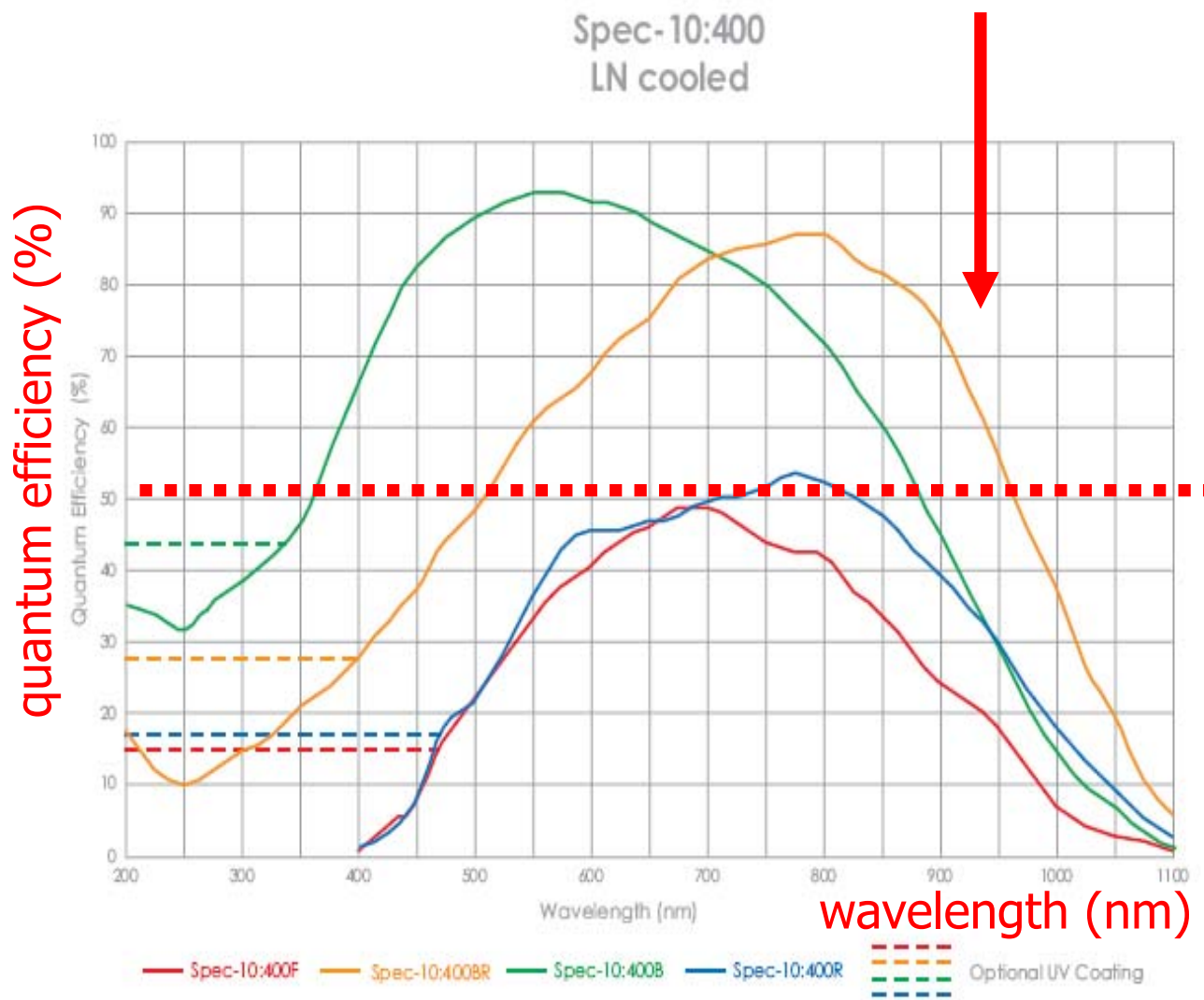


mirror lens

Spec10:400BR-LN
camera



Princeton Instruments ACTON 10:400BR-LN



q.e. high at 935 nm

50%

wavelength (nm)

LN2 cooled: **<1 e/pix/hour dark noise !!!**
used 100 kHz readout rate



CCD array: characterization

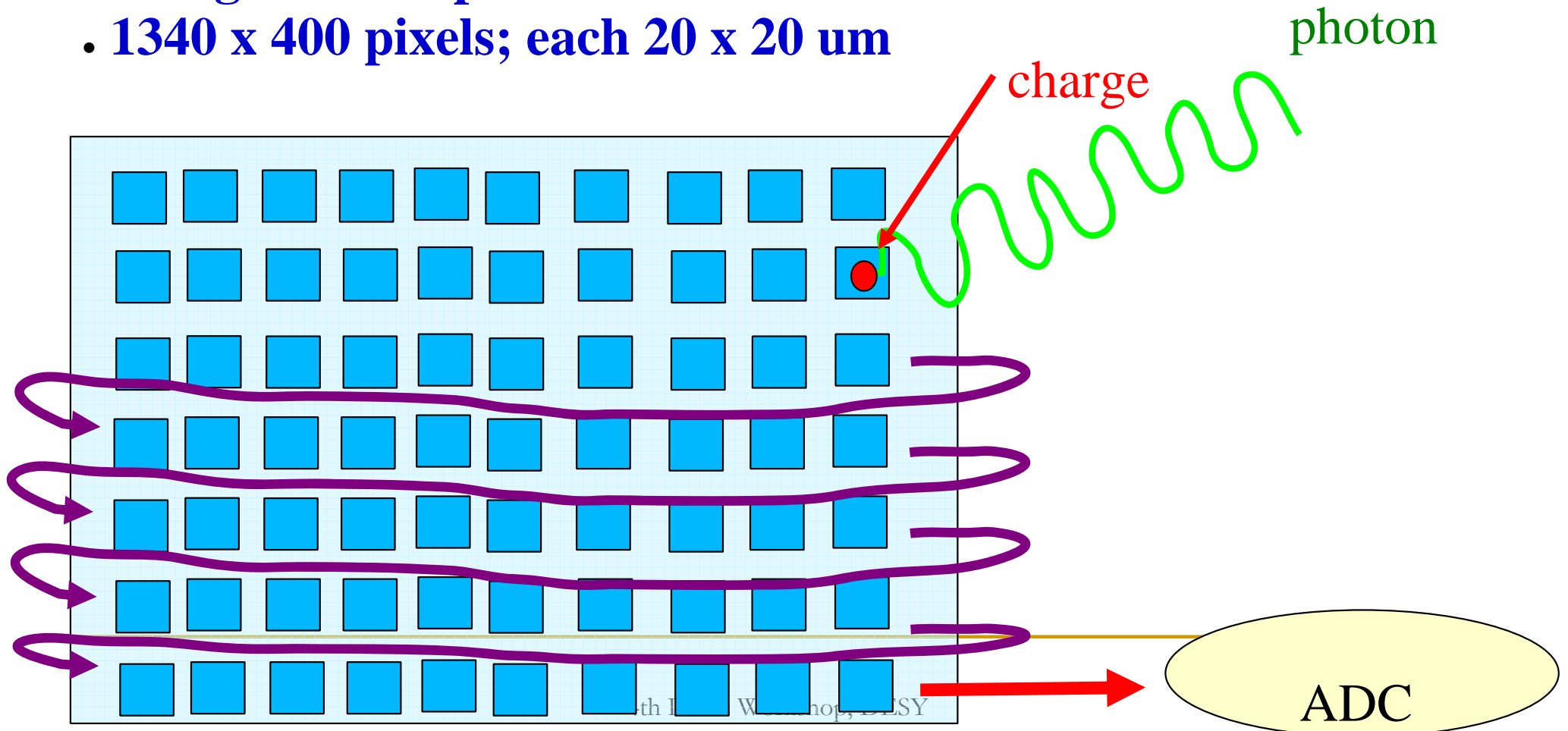
- used calibrated photon source to characterize LIPSS camera (efficiency, uniformity, . . .)
- spent past year further characterizing backgrounds and noise
 - **pixel size:** **20 μ x 20 μ**
 - **array size** **1340 x 400**
 - **gain:** **1.0 ADU/e**
 - **well depth:** **~250 ke**
 - **dynamic range** **16 bits**

backgrounds

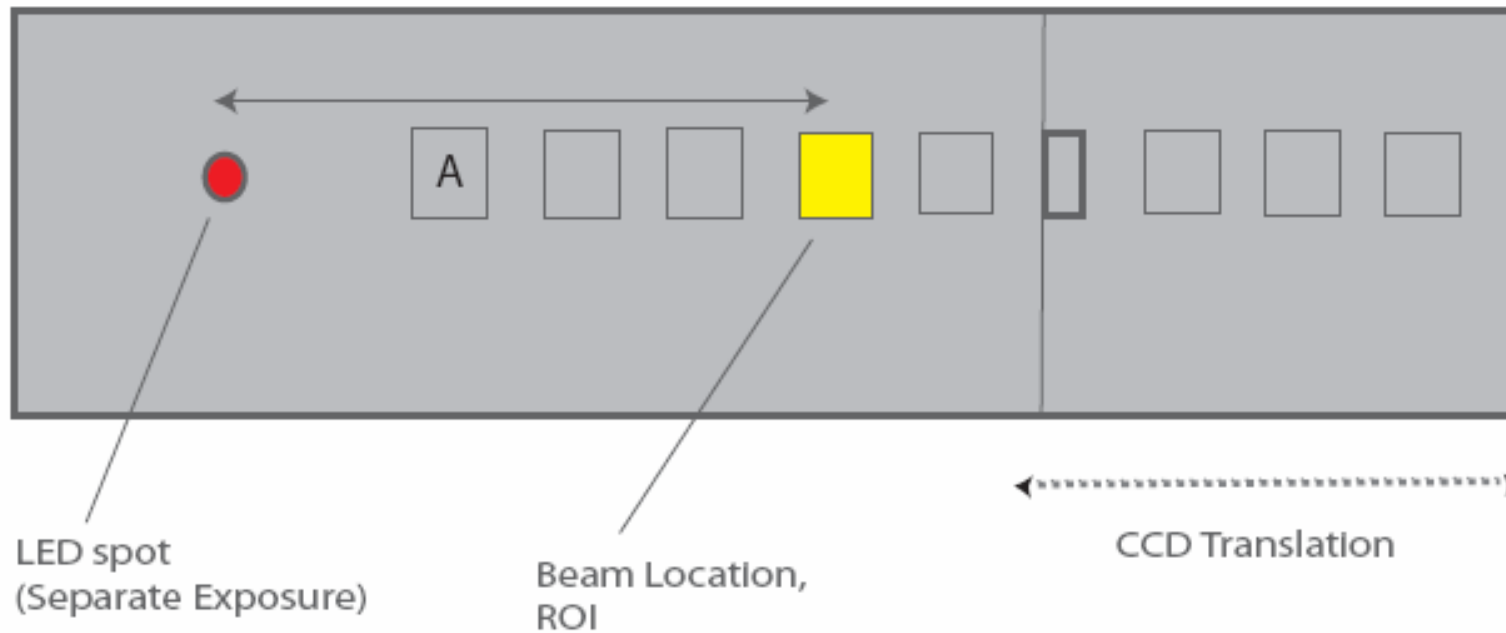
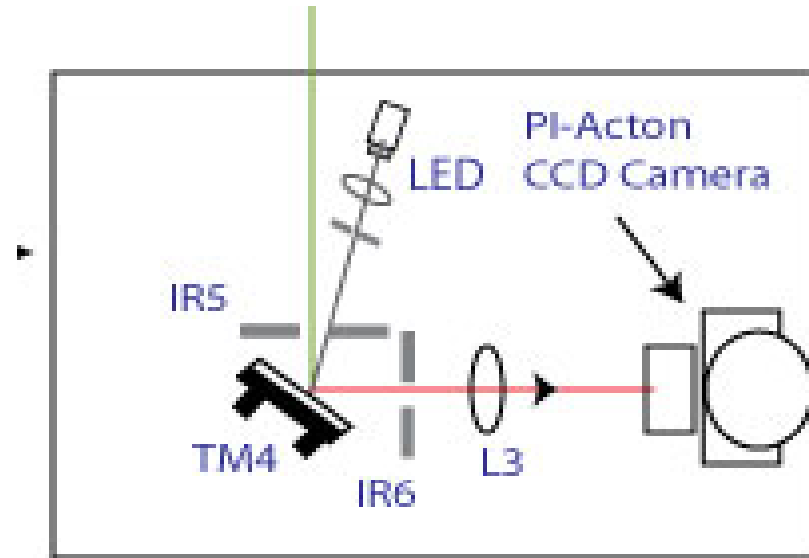
- **thermal noise**
 - **< 1 count/hour/pixel at -120°C**
- **read noise**
 - **2.7 counts per read (every 2 hours)**
- **stray light**
 - **< 1 count/hour/pixel**
- **cosmic rays in vacuum pipe gas**
 - **negligible ($\sim 10^{-6}$ Torr vacuum)**
- **cosmic rays striking CCD array**
 - **easy to identify and discard**
- **radiation from FEL**
 - **negligible**
- **...**

Piacton 400BR-LN CCD camera

- good position resolution
- no time resolution
- low dark current (cooled with LN₂)
- cosmic rays and background radiation induced events
- charge in each pixel converted to an ADU count
- 1340 x 400 pixels; each 20 x 20 um



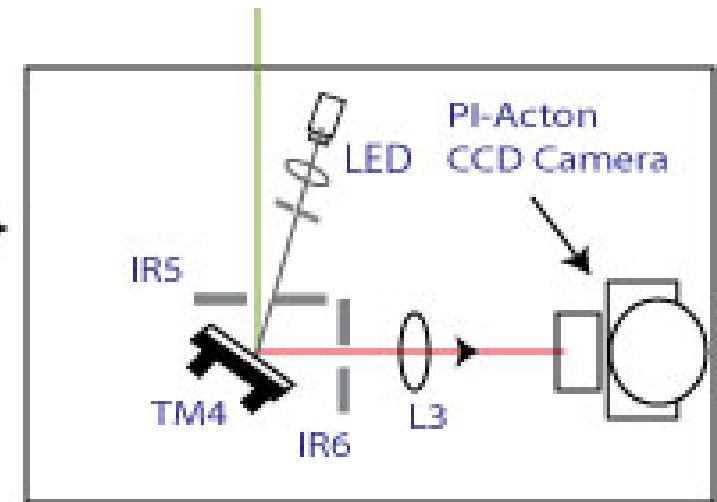
data taking . . .



increase S/N: focussing light

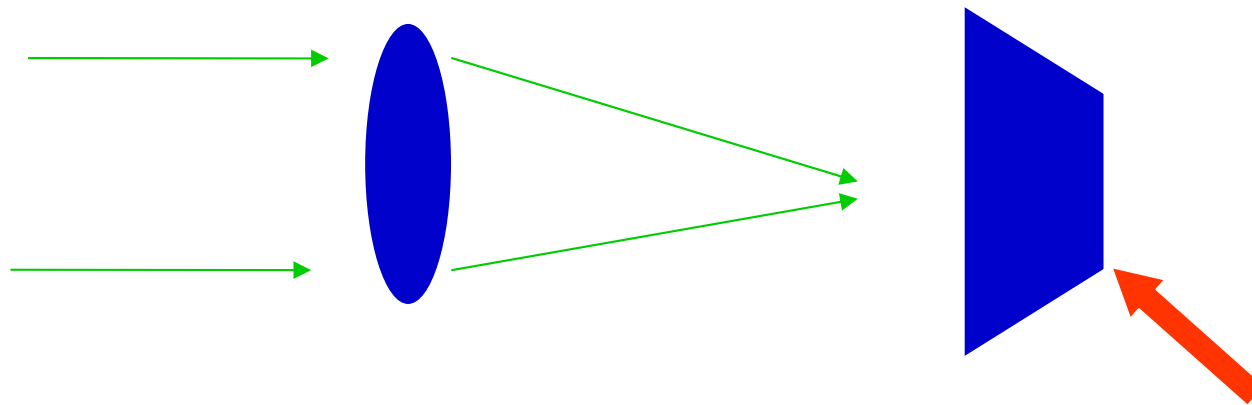
tested with FEL in alignment mode

tested with HeNe laser



focusing
lens

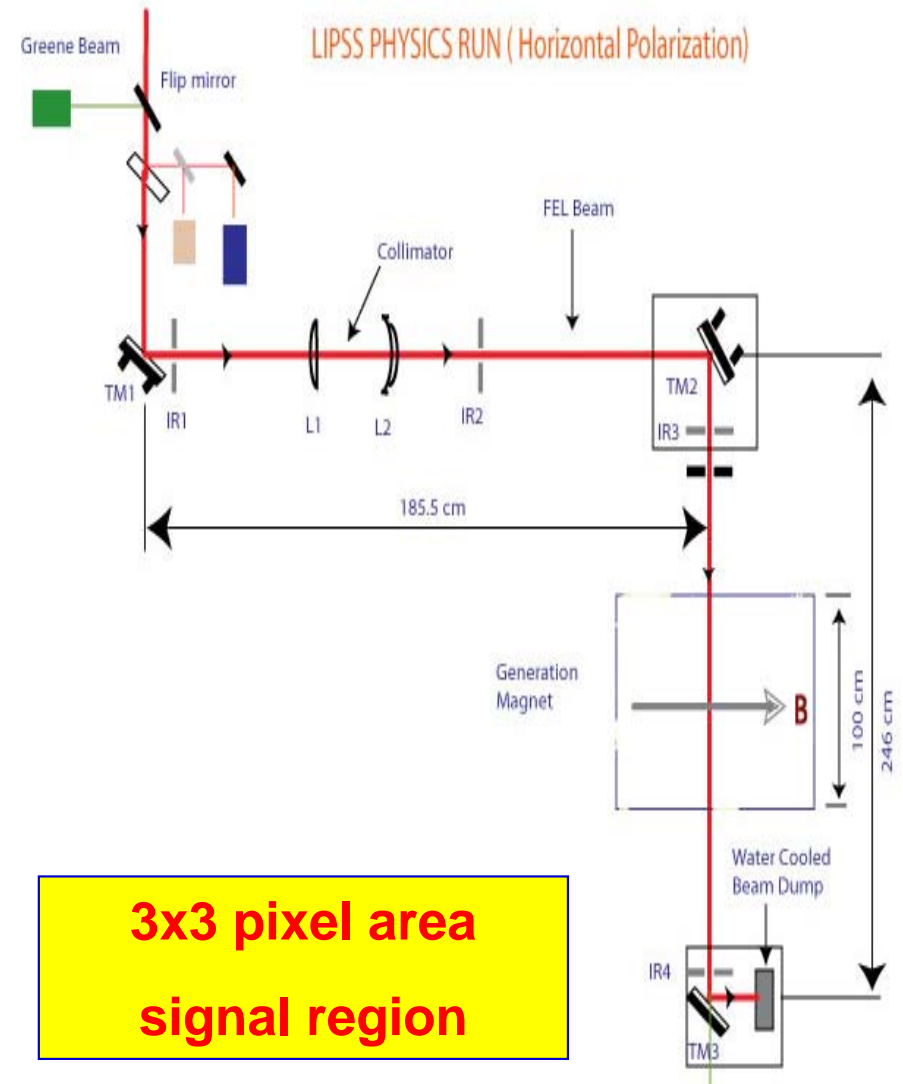
pixel array



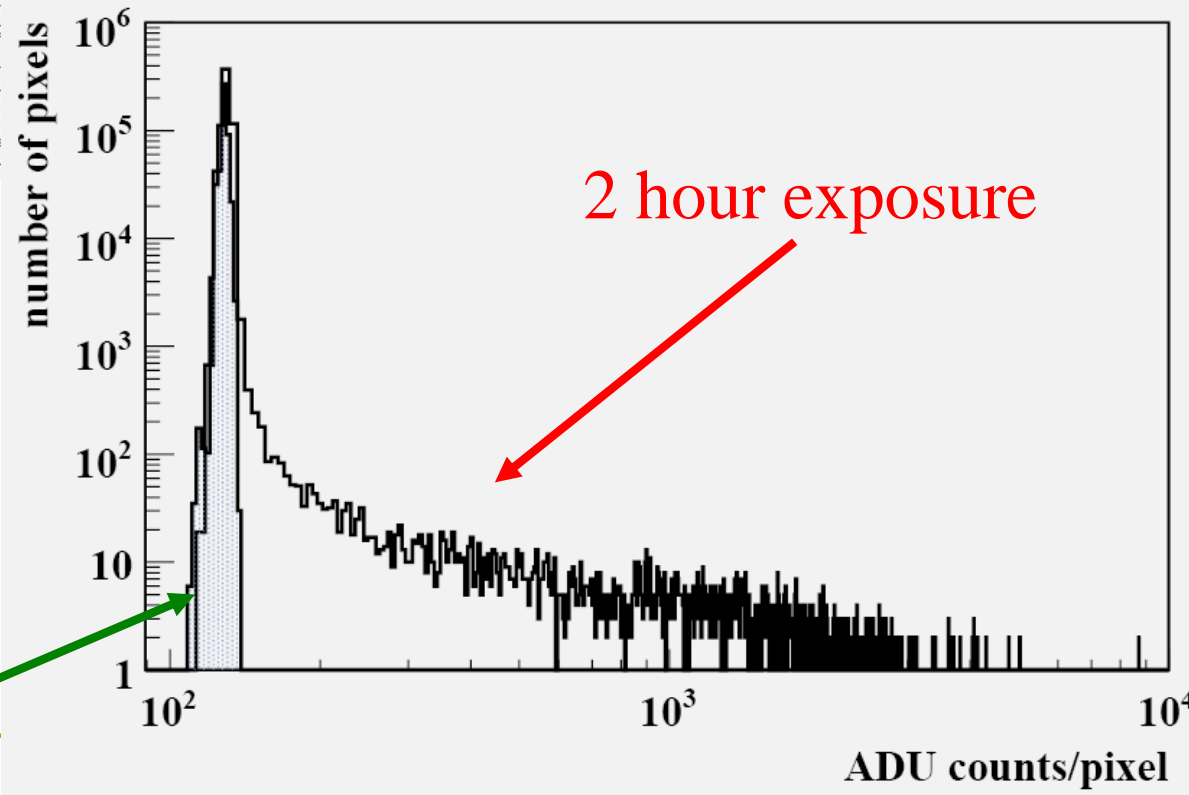
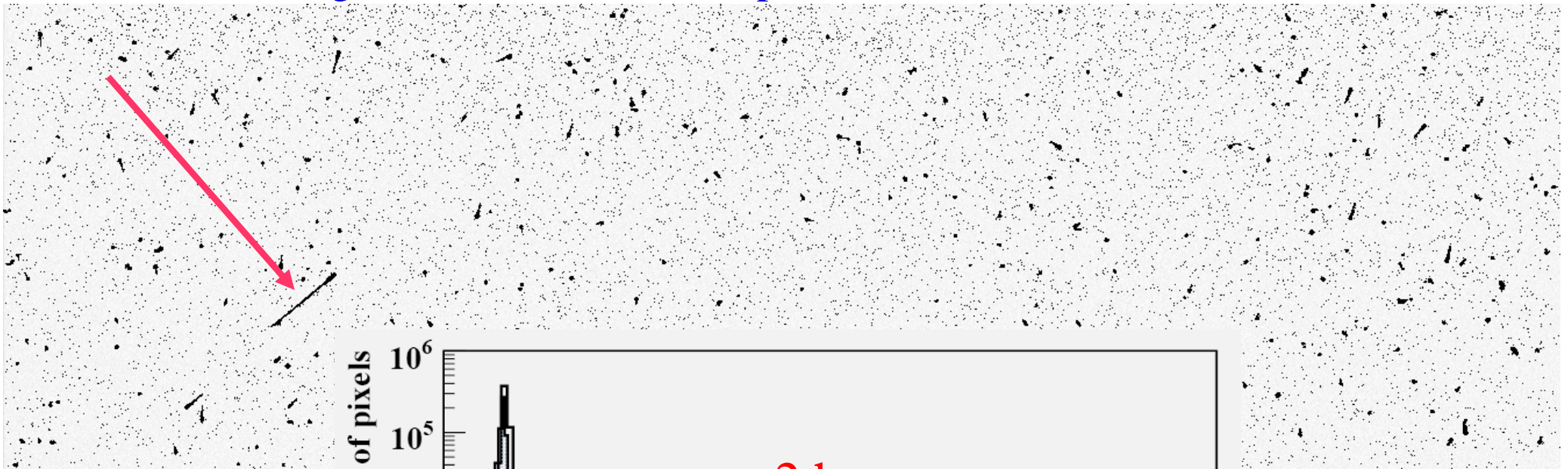
focussed to < 10
 μm spot size on
pixel

pointing error

- light rays parallel to axis are focussed onto single pixel (spot size $< 10 \mu\text{m}$)
- light rays not parallel to beam line axis can be focused onto different pixel (spherical aberration)
- controlled by keeping laser fixed at TMs
- studied this behavior during past year (with FEL and HeNe)



typical time exposure; the dark spots are due to backgrounds
(cosmic rays, stray light, thermal, . . .)
long (several hour) hour exposure #1354 from 3/2/07:

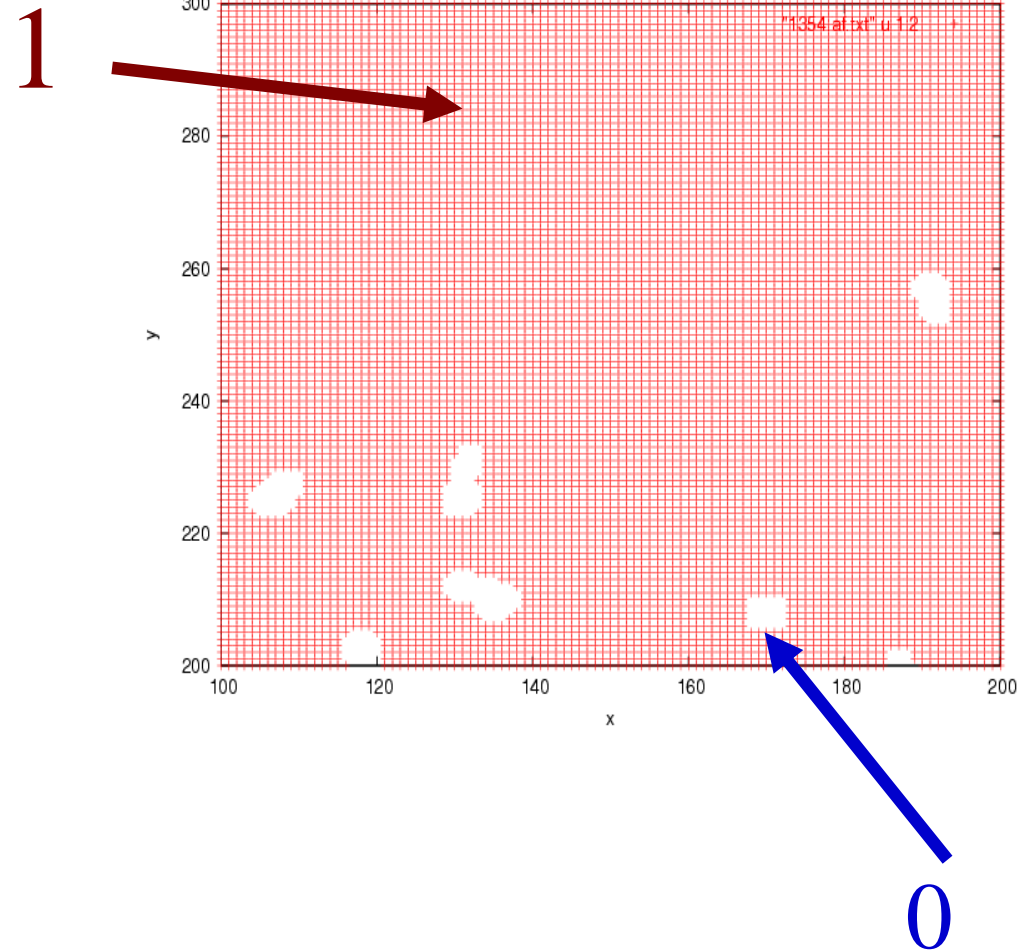
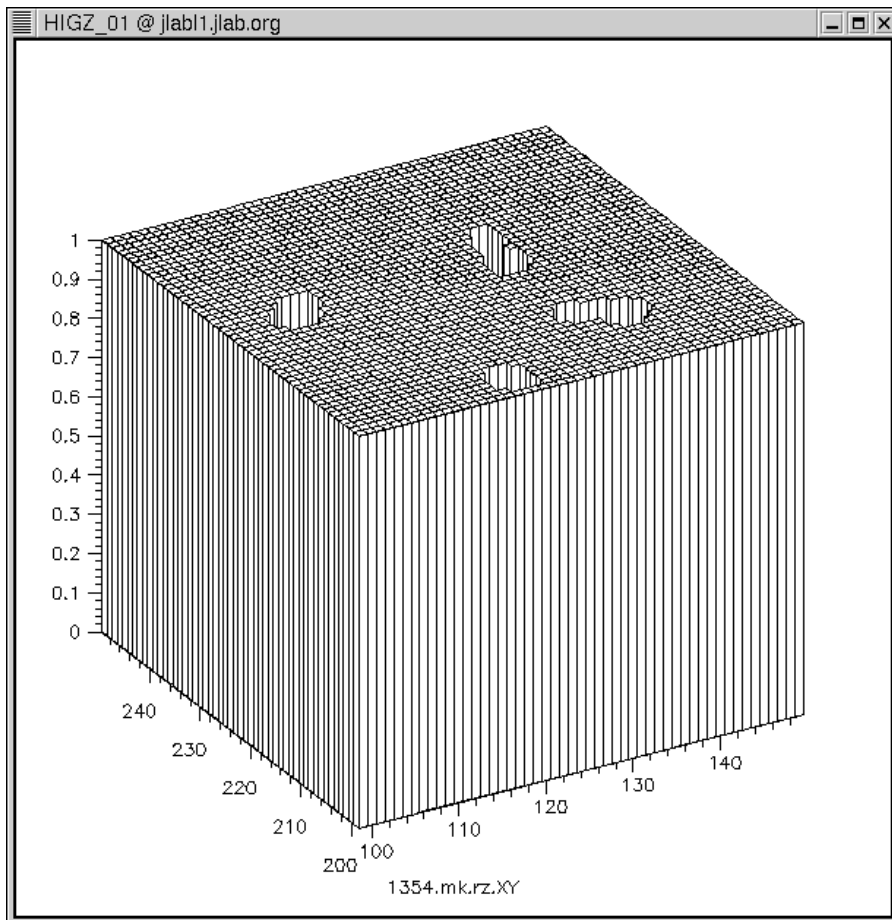


0 sec.
bias run

to remove these “anomalies”, a logical mask is created...

*estimate σ based on 5-95% cut;
then cut on $10 \times \sigma$*

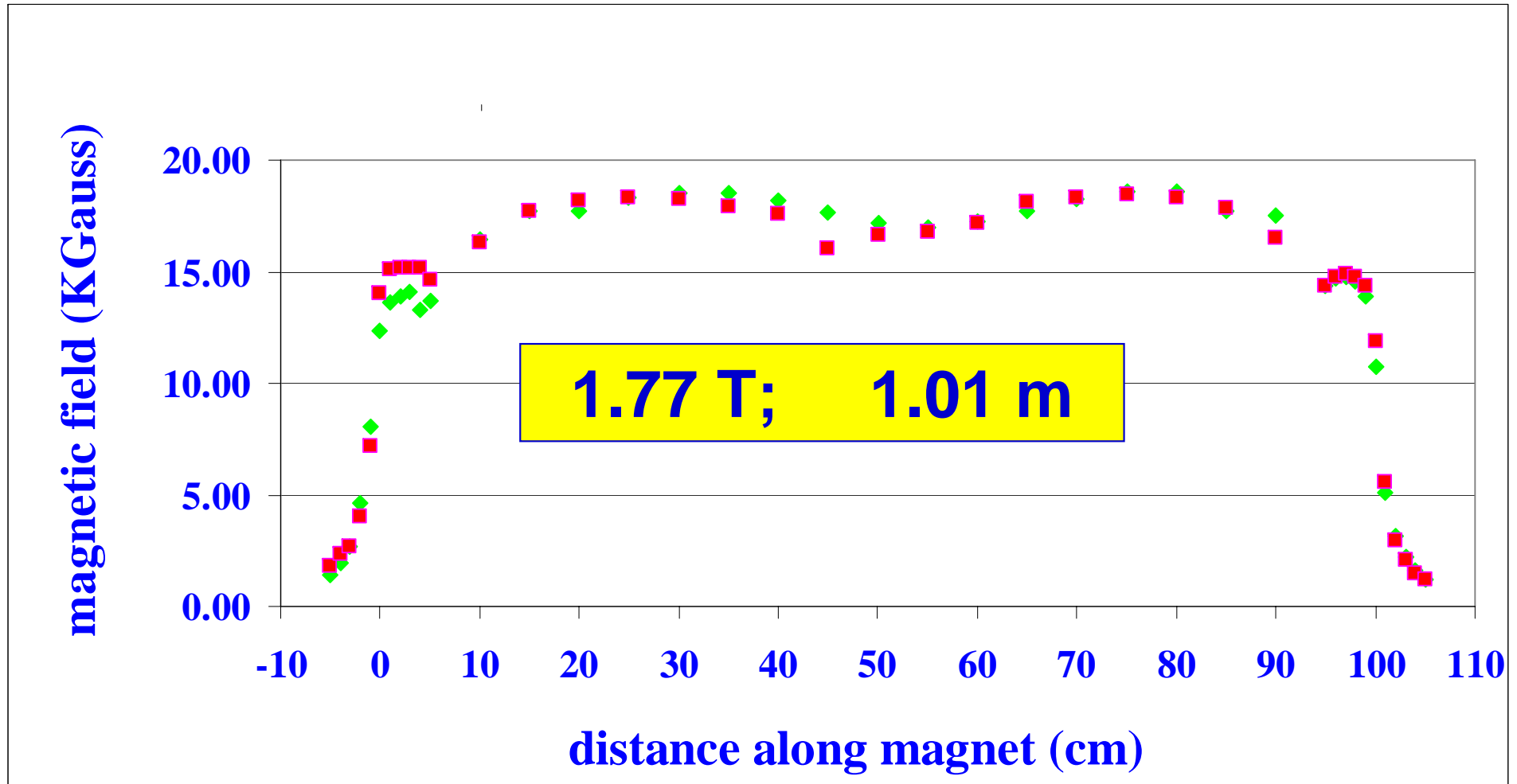
also mark off 1 pixel neighborhood



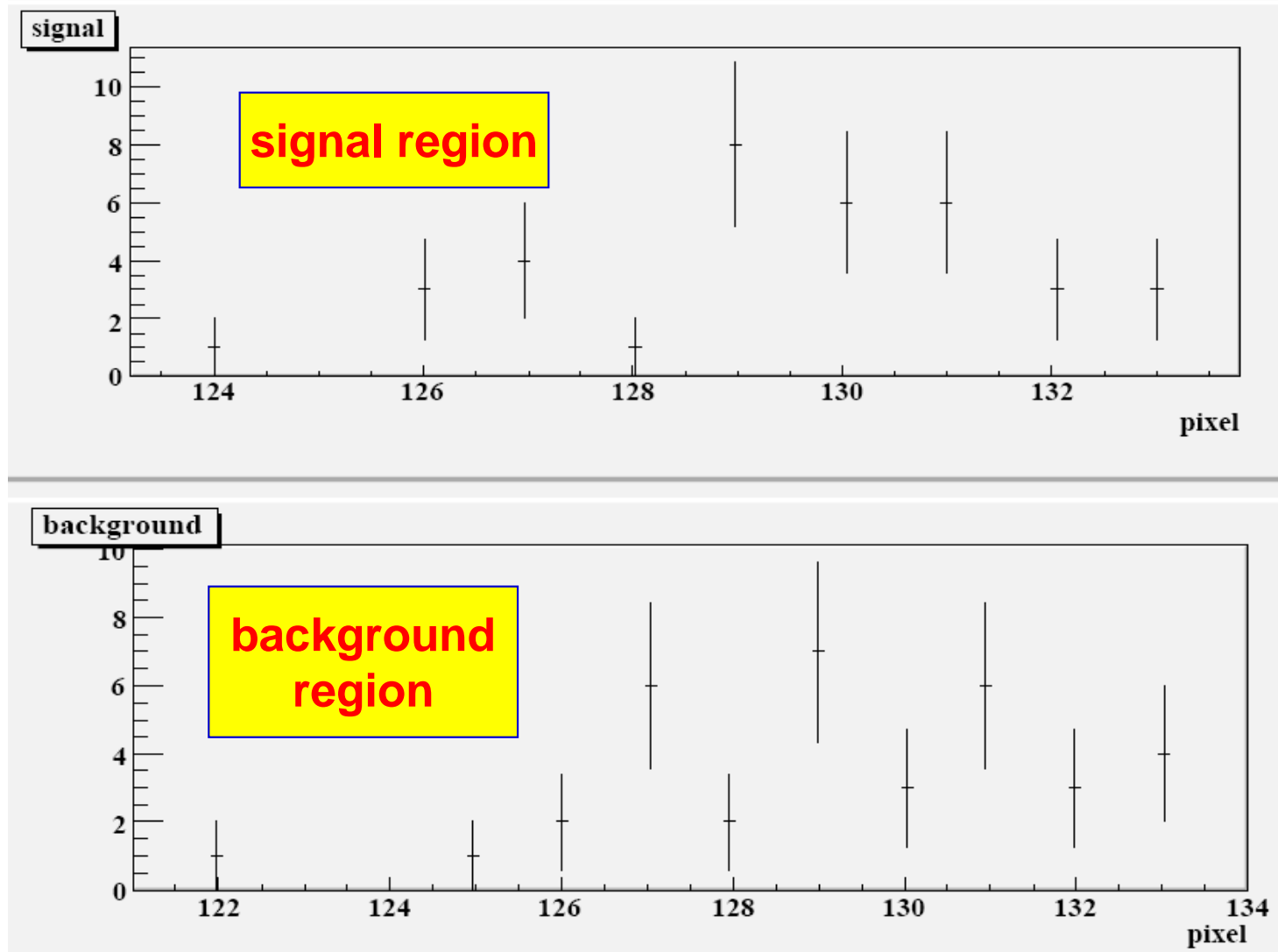
data preparation

- discarded all runs with CR hit within 100 pixels of signal region in any direction
- before each series of runs:
 - shutter open and closed
 - FEL lasing and not lasing
 - magnetic field on and off
 - room lights on and off
 - temp dependence of noise in camera
 - . . .

magnetic field strength and length



data analysis . . .



Spring '07 data

- 17 hours of data taking (over a period of one week).
- laser light polarized perpendicular to magnetic field direction (scalar coupling).
- backgrounds well characterized over the past two years.
 - **large number of background pixels (several million)**
- laser light focussing and pointing error well characterized over the past two years.
 - **focused to 3x3 pixel signal area**

parameters for initial LIPSS run (2007)

- B-field: 1.77 T
- magnet length: 1.01 m
- IR FEL power 0.18 kW
- IR FEL wavelength 935 nm (1.33 eV)
- quantum efficiency 0.45
- linear polarization 100%
- acceptance 100%
- experimental efficiency ~ 90%

results

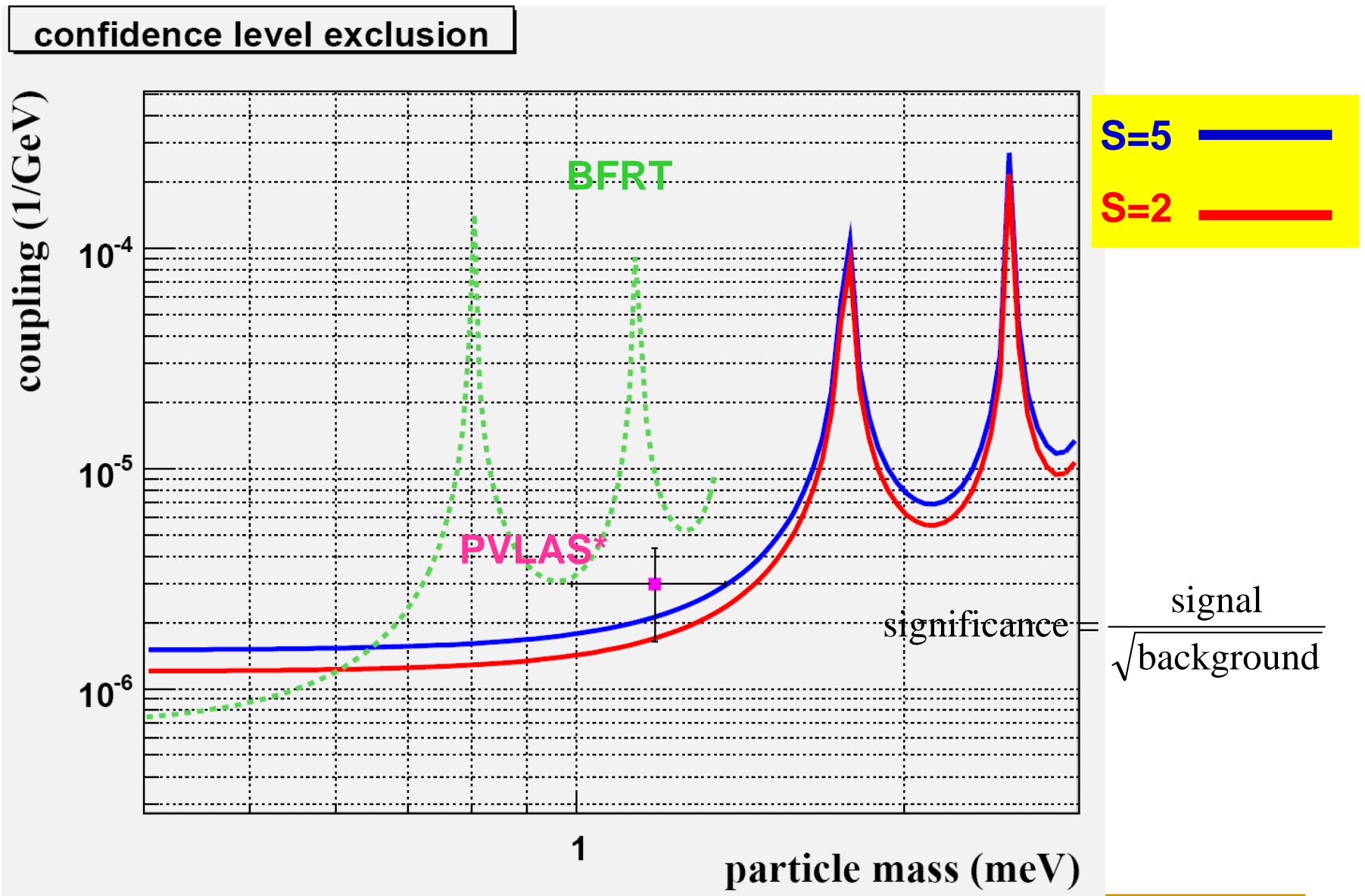
$$Y = n P_1 P_2 \varepsilon (\Delta\Omega/\Omega) \quad \text{yield (\#/s)}$$

- n = photon flux (#/s)
- P_1 (P_2) = production (regeneration) probability
- ε = detection efficiency
- $\Delta\Omega/\Omega$ = solid angle

$$S \equiv \frac{R_s \times t}{\sqrt{R_b \times t}} \quad S \geq 5$$

significance greater than or equal to five for discovery

significance of LIPSS result



* now disclaimed

LIPSS at Jefferson Lab

- no evidence for scalar coupling at this level
- will resume running at Jefferson Lab in Fall 08
 - upgraded FEL
 - upgraded LIPSS equipment
 - will run in pseudoscalar mode
- developing a new experiment using microwave cavities (see P. Slocum's talk this conference)
 - 34 GHz source
 - resonant cavities

LIPSS collaboration

O.K. Baker*, M. Minarni †, P. Slocum
Yale University

A. Afanasev, R. Ramdon
Hampton University

K. Beard#, G. Biallas, J. Boyce, M. Shinn
Jefferson Lab

**Spokesman †now at Riau Univ, Indonesia #now at Muons Inc, Batavia*

acknowledgements

- **FEL Division:** F. Dylla, G. Neil, G. Williams, R. Walker, D. Douglas, S. Benson, K. Jordan, C. Hernandez-Garcia, J. Gubeli
- **Hampton Univ:** M.C. Long, K. McFarlane

funding

- **Office of Naval Research**
- **Yale University**